

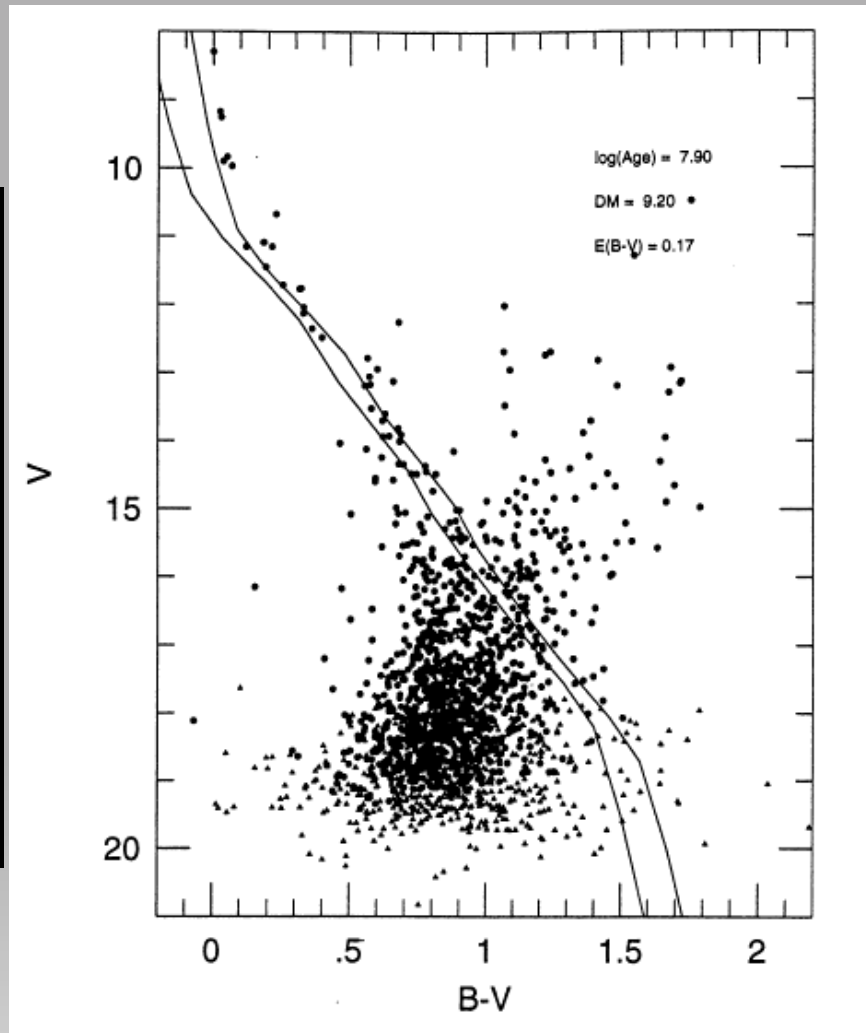
# Video Camera Spectral Sensitivity

Hristo Pavlov  
IOTA, WSAAG, Tangra Observatory

Lucky Star Kick-Off Meeting Workshop  
Paris, 18-19 Apr 2016



**NGC 6716**



- 14" LX200 ACF + F/3.3 FR
- Flea3, WAT120N+, WAT920H, PC164EX2
- AVer Grabber, VDub, HuffYuv & MPEG4
- SLOAN Photometric Filters
- ADVS

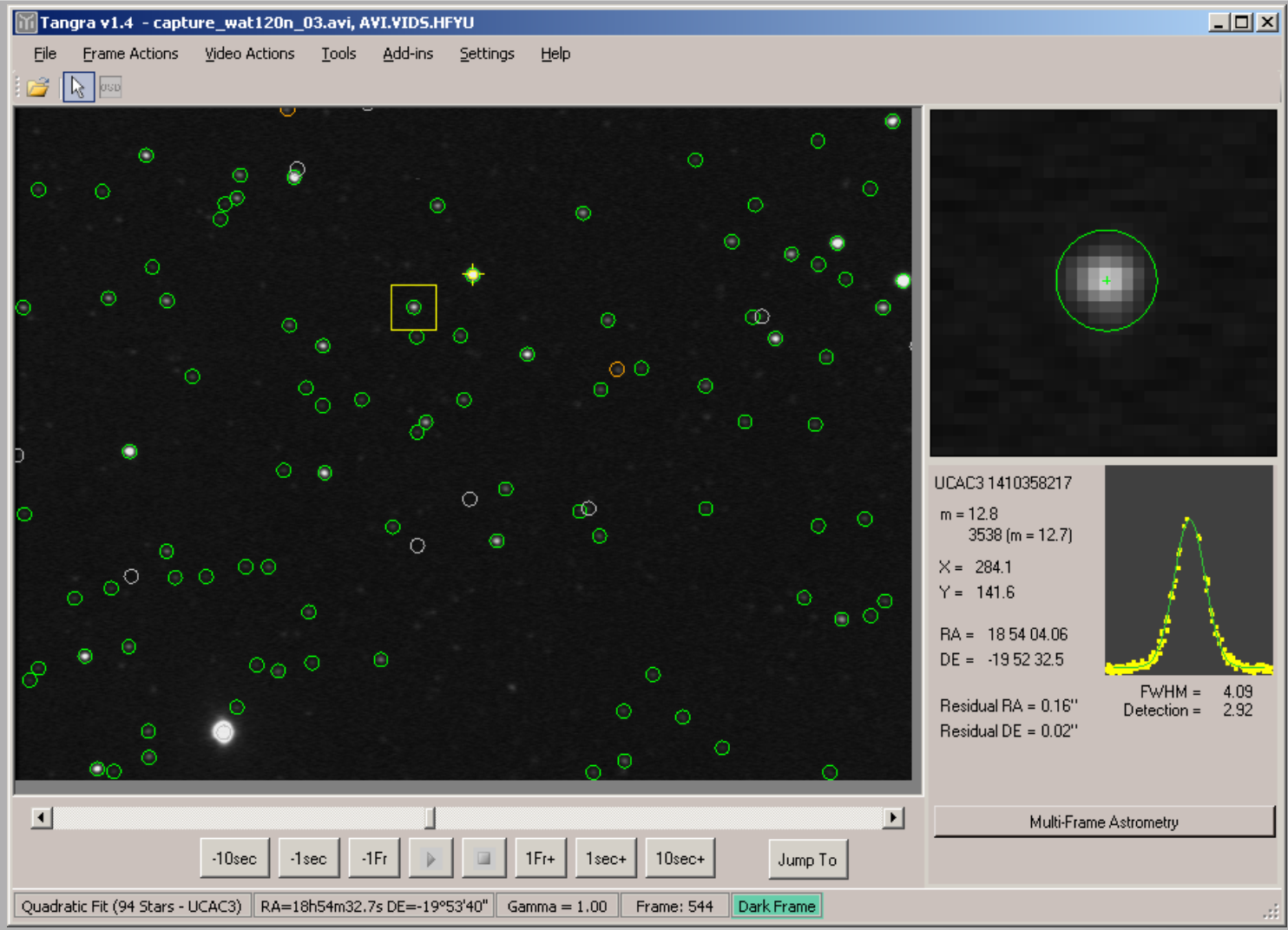


**Equipment**



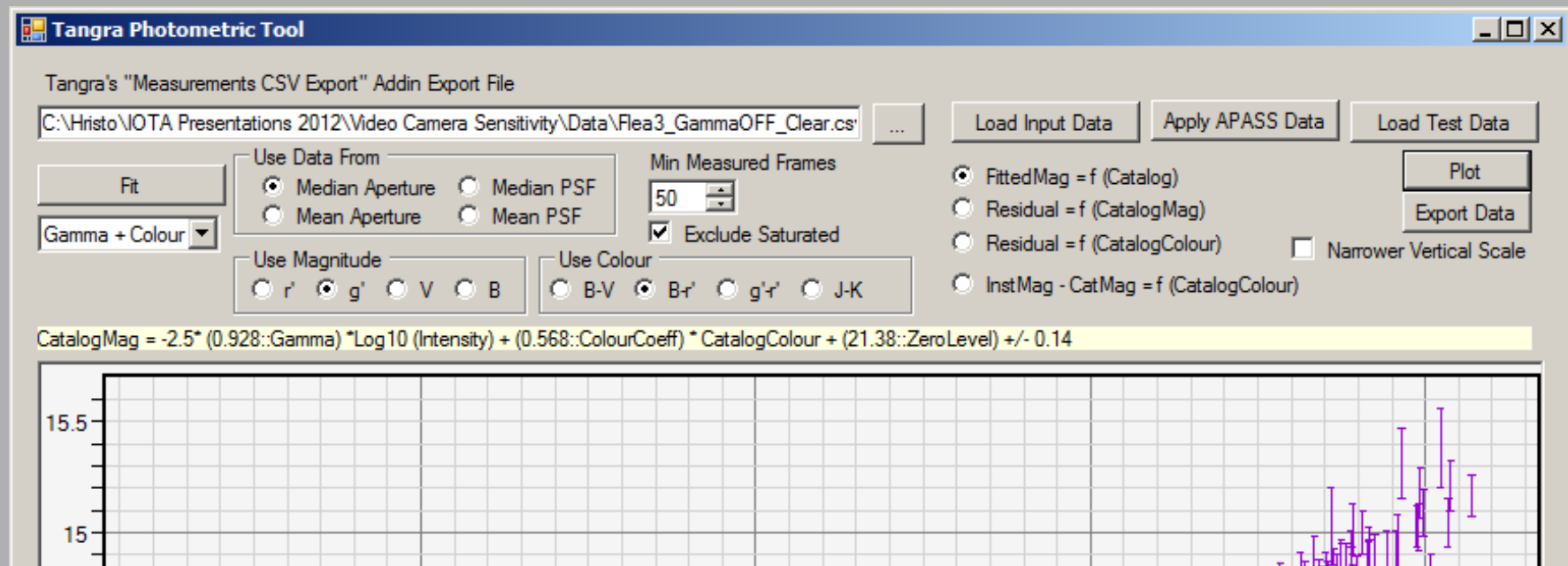
| Date   | Camera   | Start UT | End UT | Filter | Zenith Angle | AtmExt Max $\Delta m$ |
|--------|----------|----------|--------|--------|--------------|-----------------------|
| 03 Sep | WAT120N  | 11:10    | 11:24  | Clear  | 21°-24°      | 0.01 <sup>m</sup>     |
| 21 Sep | Flea3    | 10:53    | 11:07  | r'     | 31°-34°      | 0.01 <sup>m</sup>     |
| 21 Sep | Flea3    | 11:09    | 11:20  | r'     | 35°-37°      | 0.01 <sup>m</sup>     |
| 21 Sep | Flea3    | 11:37    | 11:48  | Clear  | 40°-43°      | 0.02 <sup>m</sup>     |
| 04 Oct | WAT120N  | 10:59    | 11:14  | Clear  | 43°-46°      | 0.02 <sup>m</sup>     |
| 04 Oct | WAT902H  | 11:20    | 11:32  | Clear  | 47°-50°      | 0.02 <sup>m</sup>     |
| 04 Oct | PC164EX2 | 11:39    | 11:45  | Clear  | 50°-53°      | 0.02 <sup>m</sup>     |

## Observing Sessions



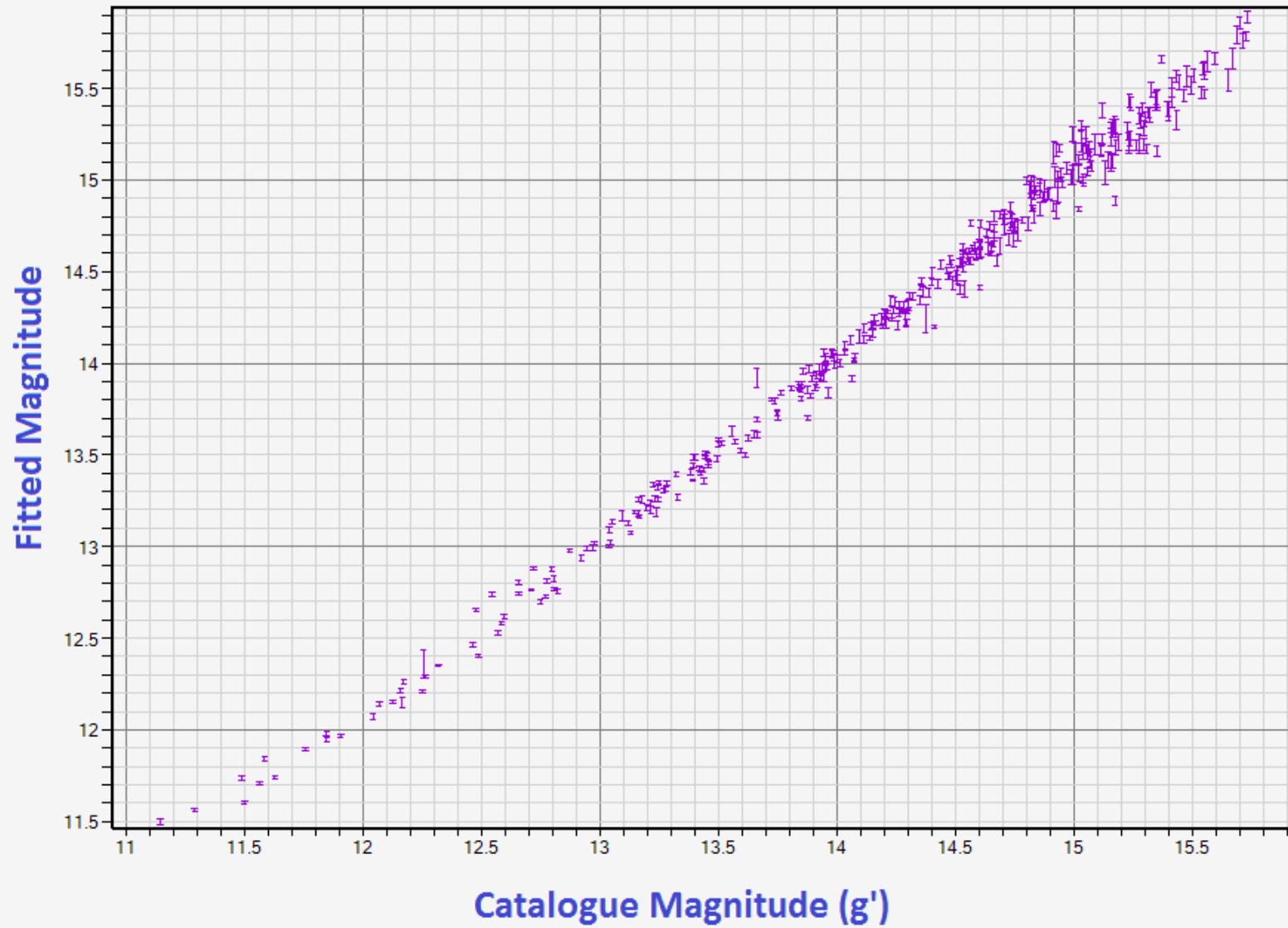
# Data Reduction

$$\text{Mag} = -2.5 * \text{NonLinerCoeff} * \text{Log10}(I) + \text{ColourCoeff} * \text{CatalogColour} + \text{ZeroPoint}$$

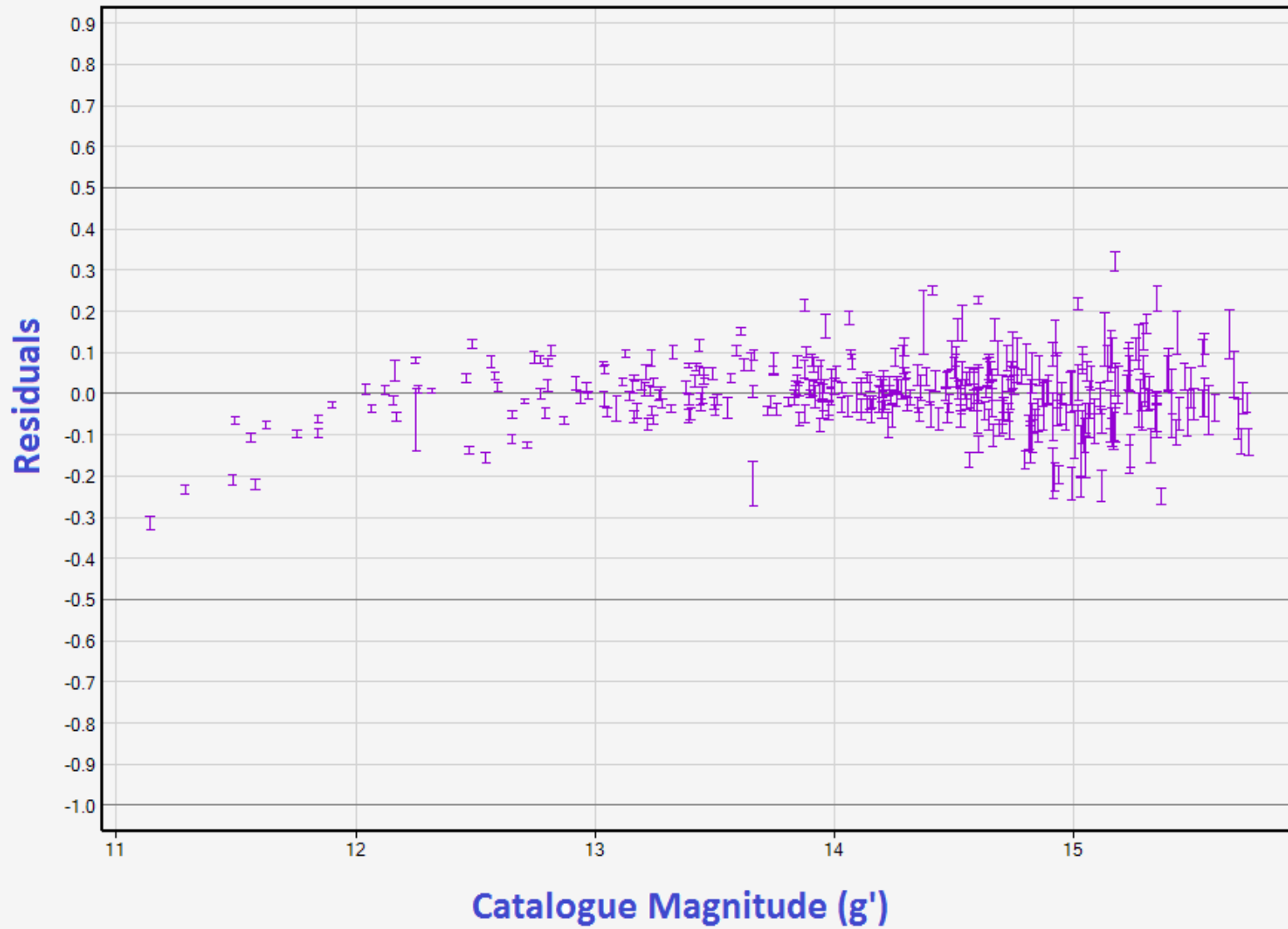


# Magnitude and Colour Fitting

# WAT120N+, Clear

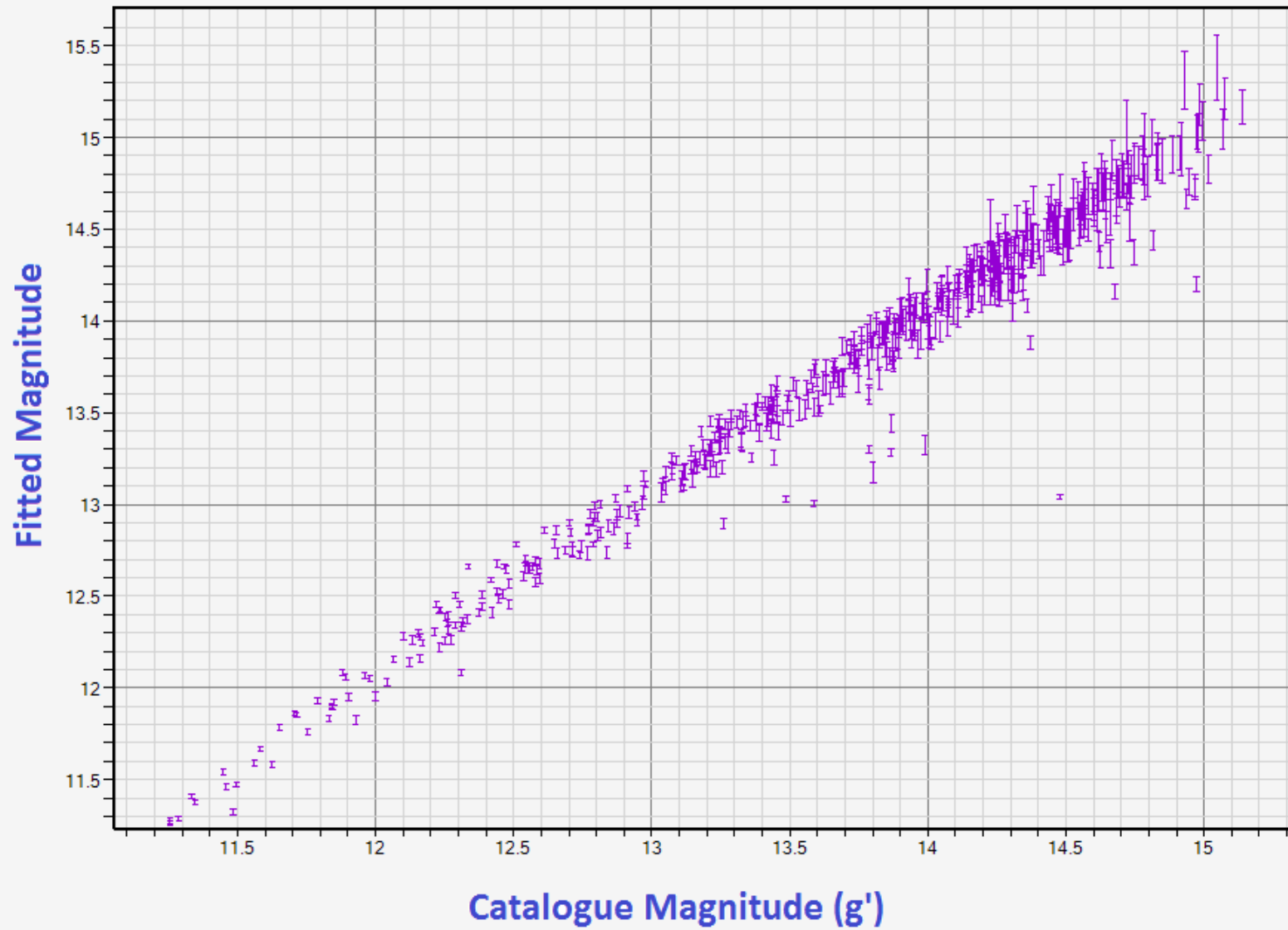


# WAT120N+, Clear

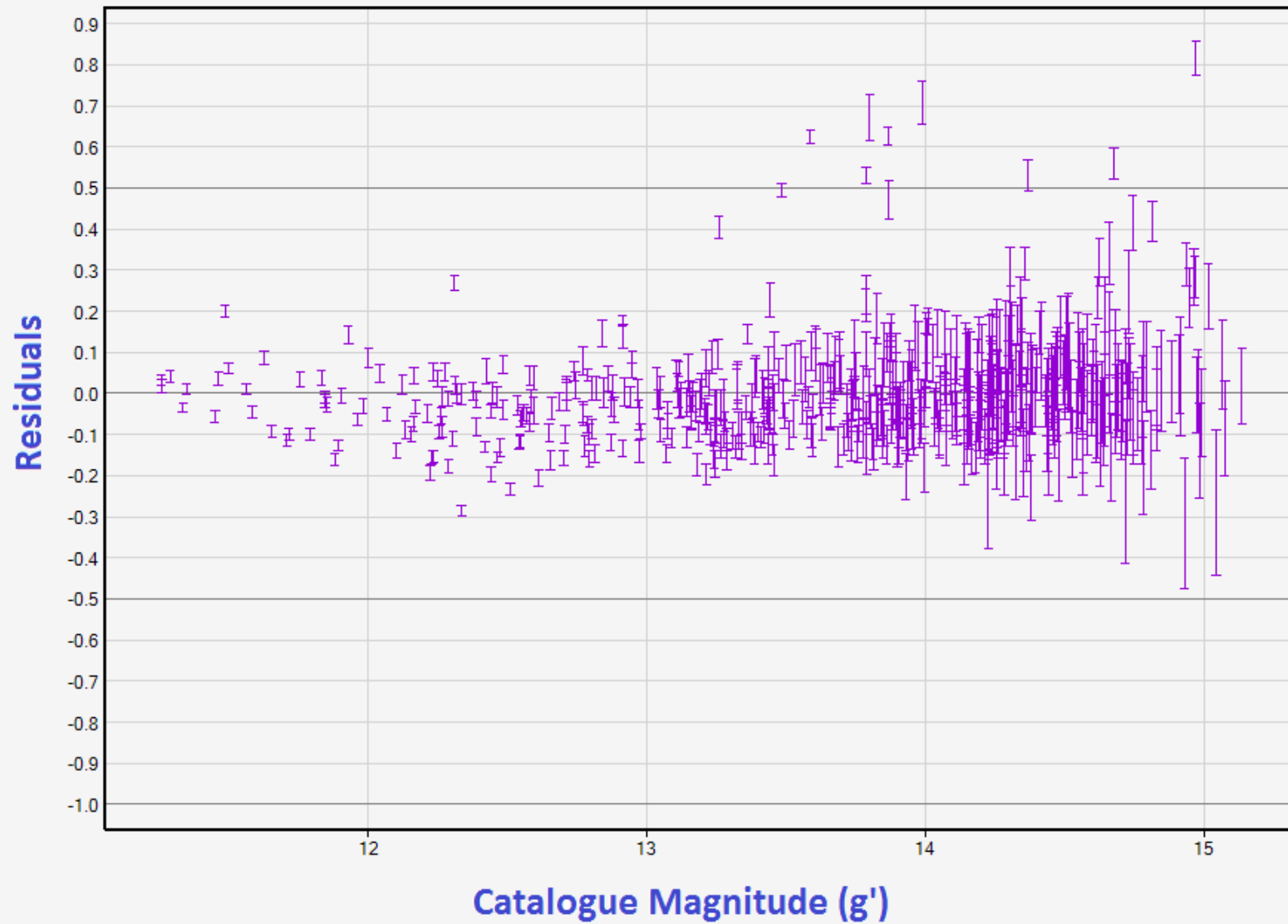




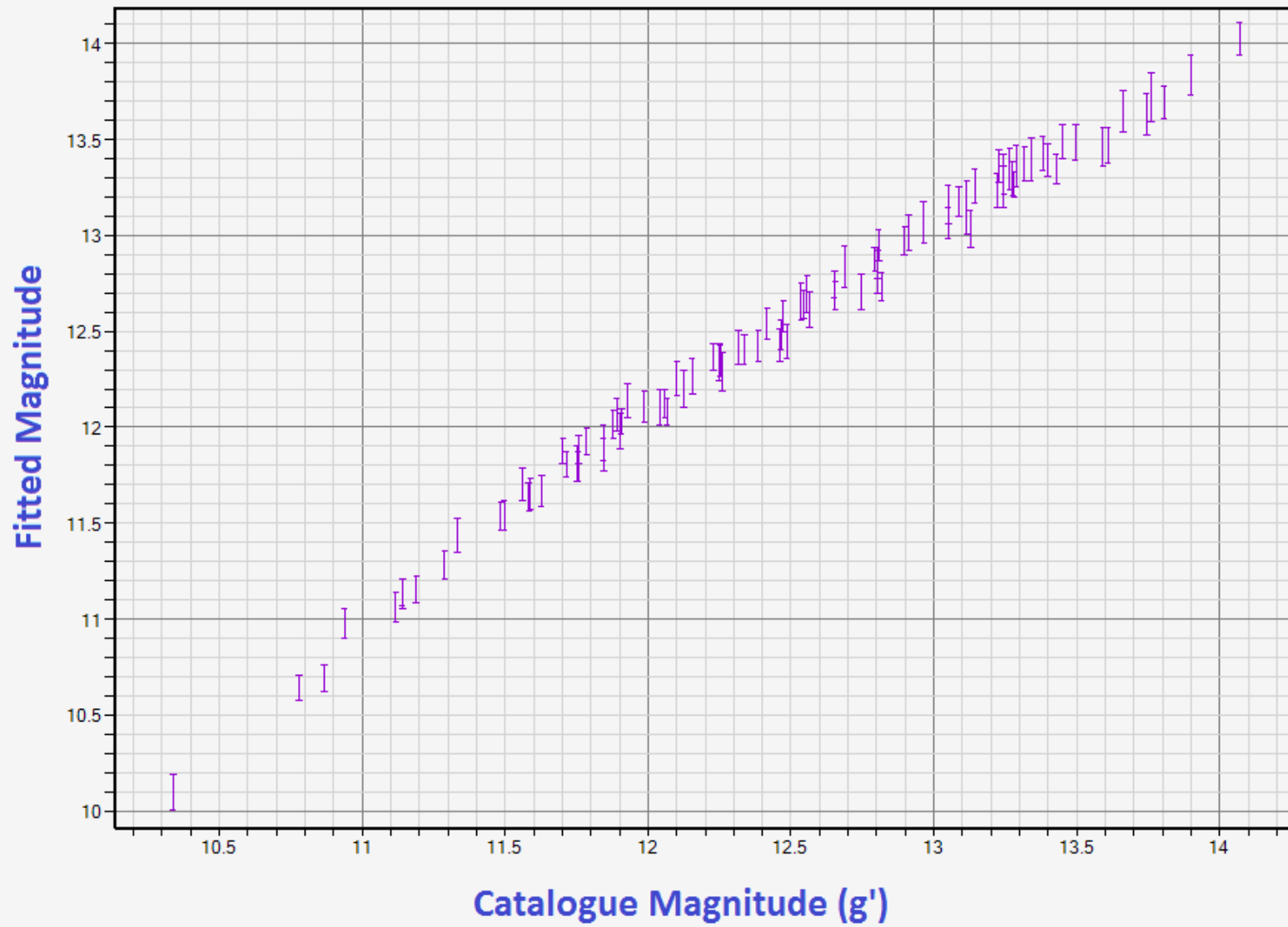
# Flea3, Clear



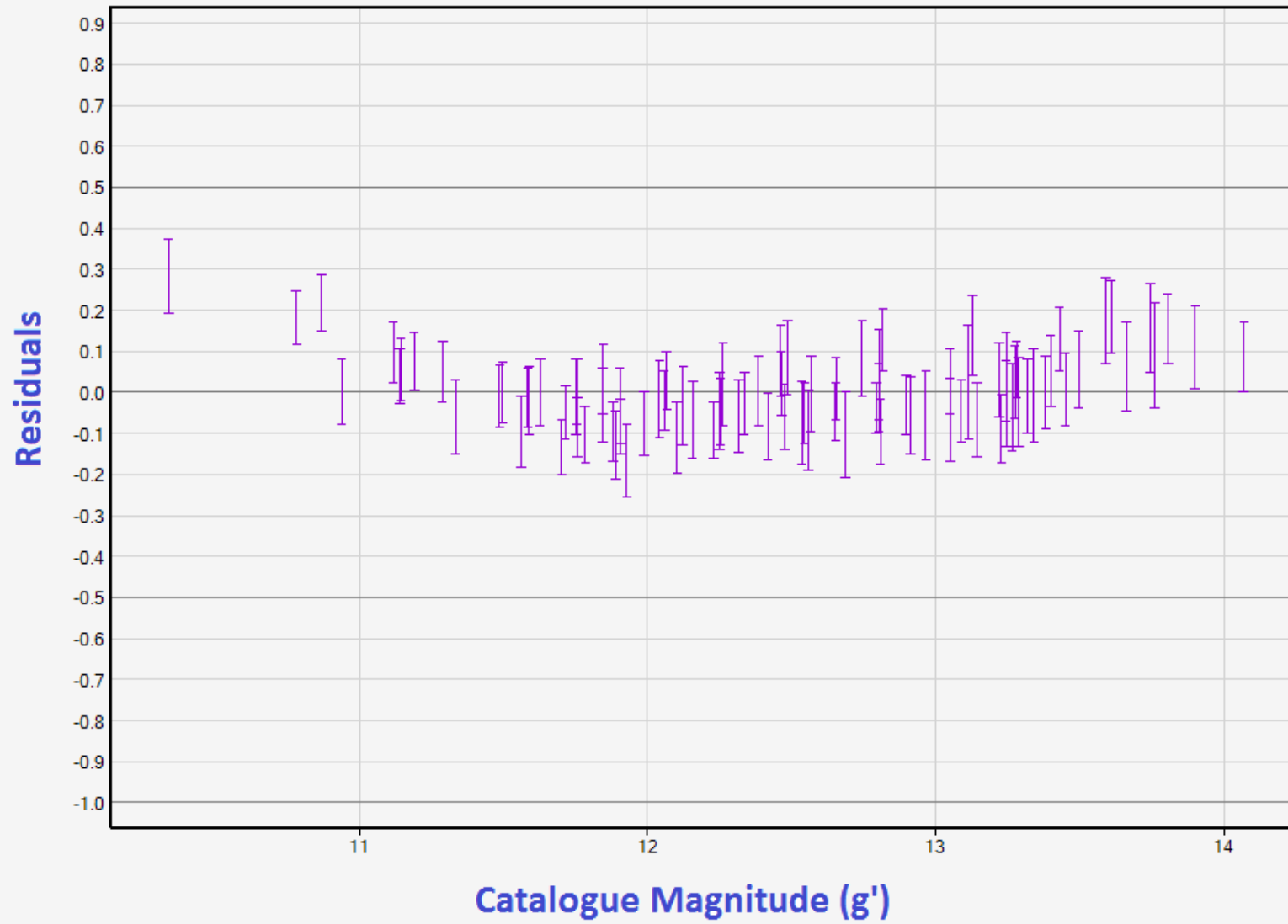
# Flea3, Clear



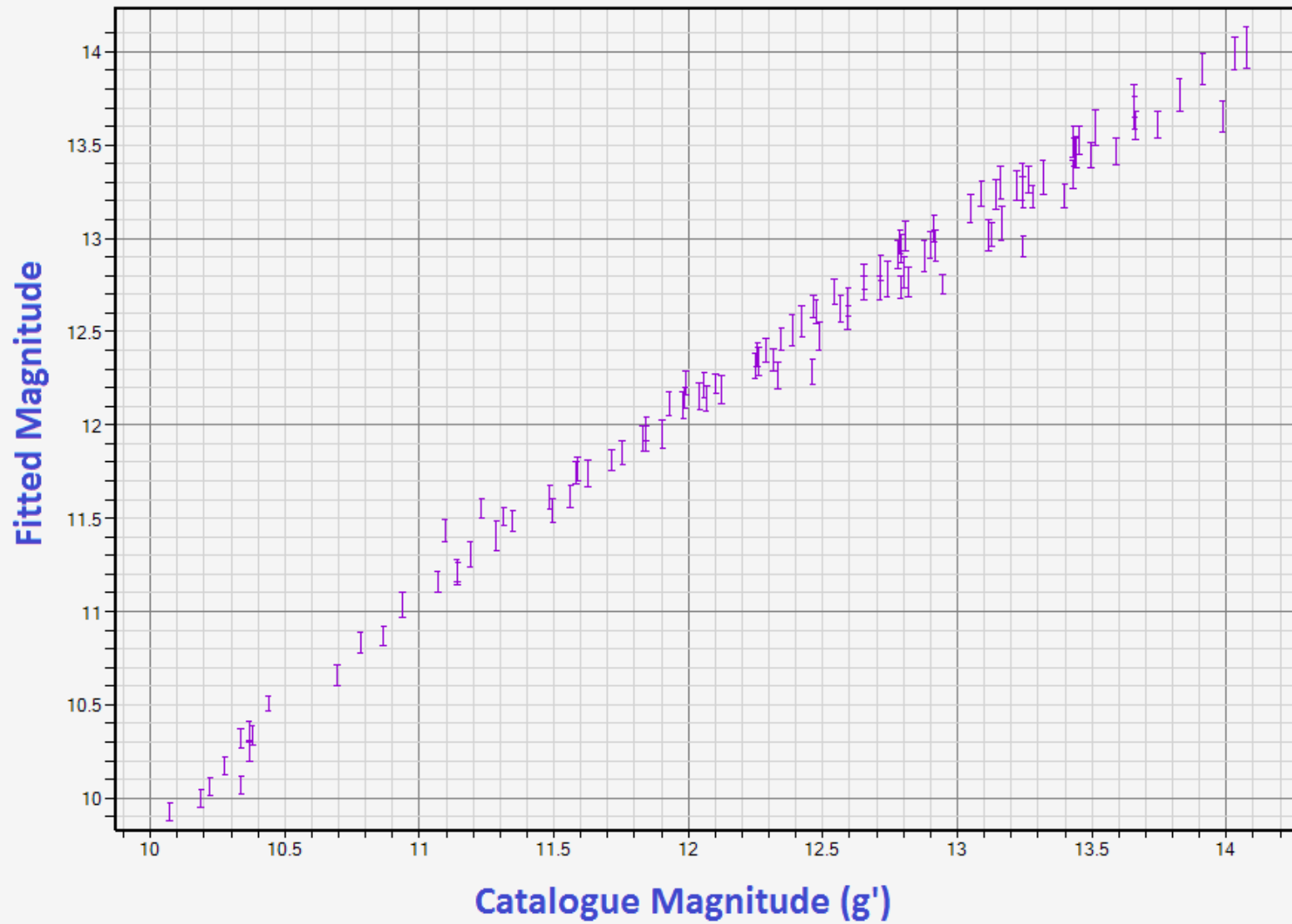
# WAT902H, Clear



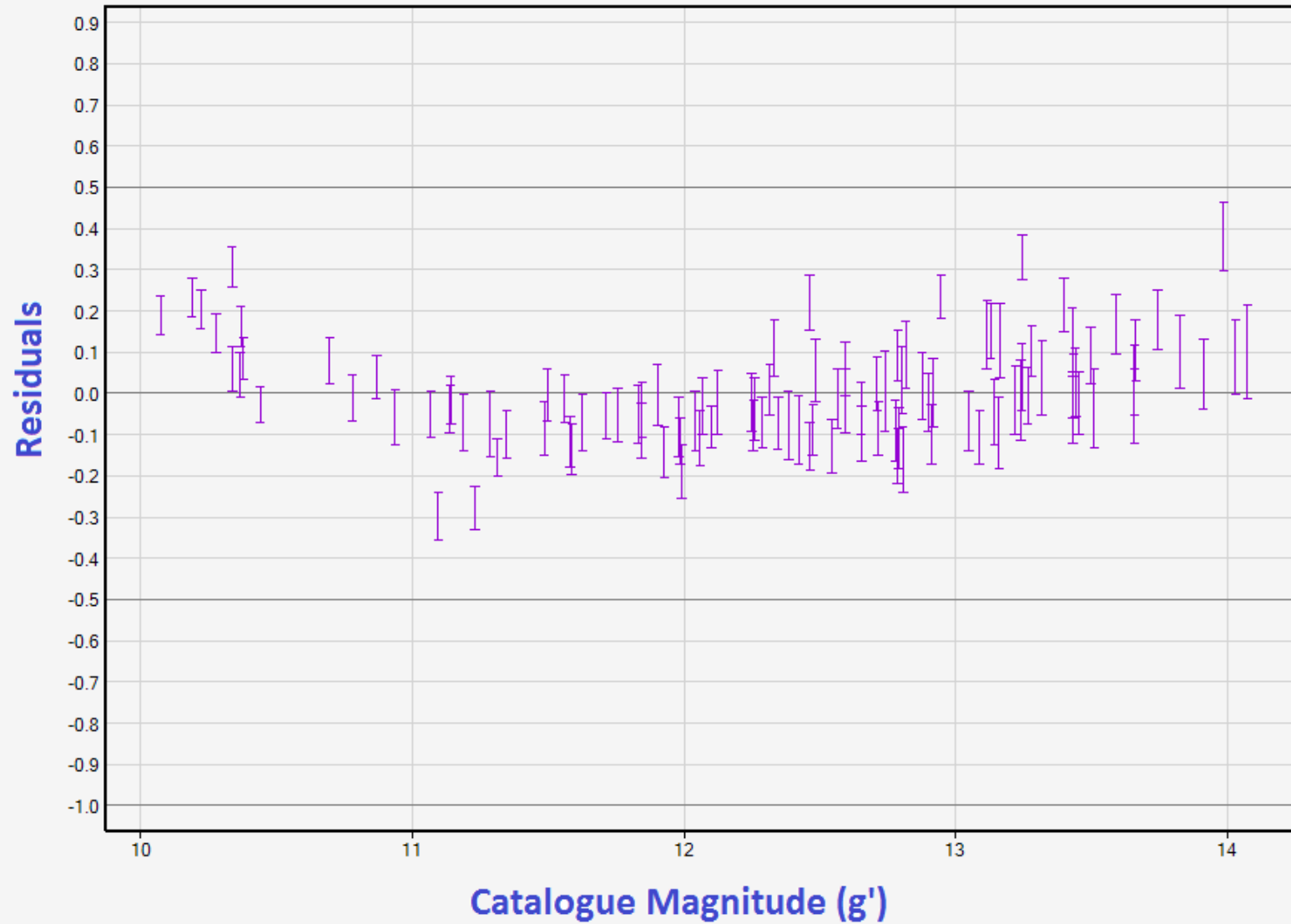
# WAT902H, Clear



# PC164-EX2



# PC164-EX2



### WAT120N+ (Gamma = OFF)

$$g' = -2.5 * 1.012 * \text{Log}_{10} (\text{Intensity}) + 0.549 * (B-r') + 21.36 \pm 0.08$$

$$r' = -2.5 * 1.017 * \text{Log}_{10} (\text{Intensity}) - 0.121 * (B-r') + 21.50 \pm 0.08$$

### WAT902H (Gamma = OFF)

$$g' = -2.5 * 1.183 * \text{Log}_{10} (\text{Intensity}) + 0.596 * (B-r') + 20.13 \pm 0.08$$

$$r' = -2.5 * 1.210 * \text{Log}_{10} (\text{Intensity}) - 0.079 * (B-r') + 20.43 \pm 0.07$$

### PC164EX2

$$g' = -2.5 * 1.034 * \text{Log}_{10} (\text{Intensity}) + 0.614 * (B-r') + 19.47 \pm 0.12$$

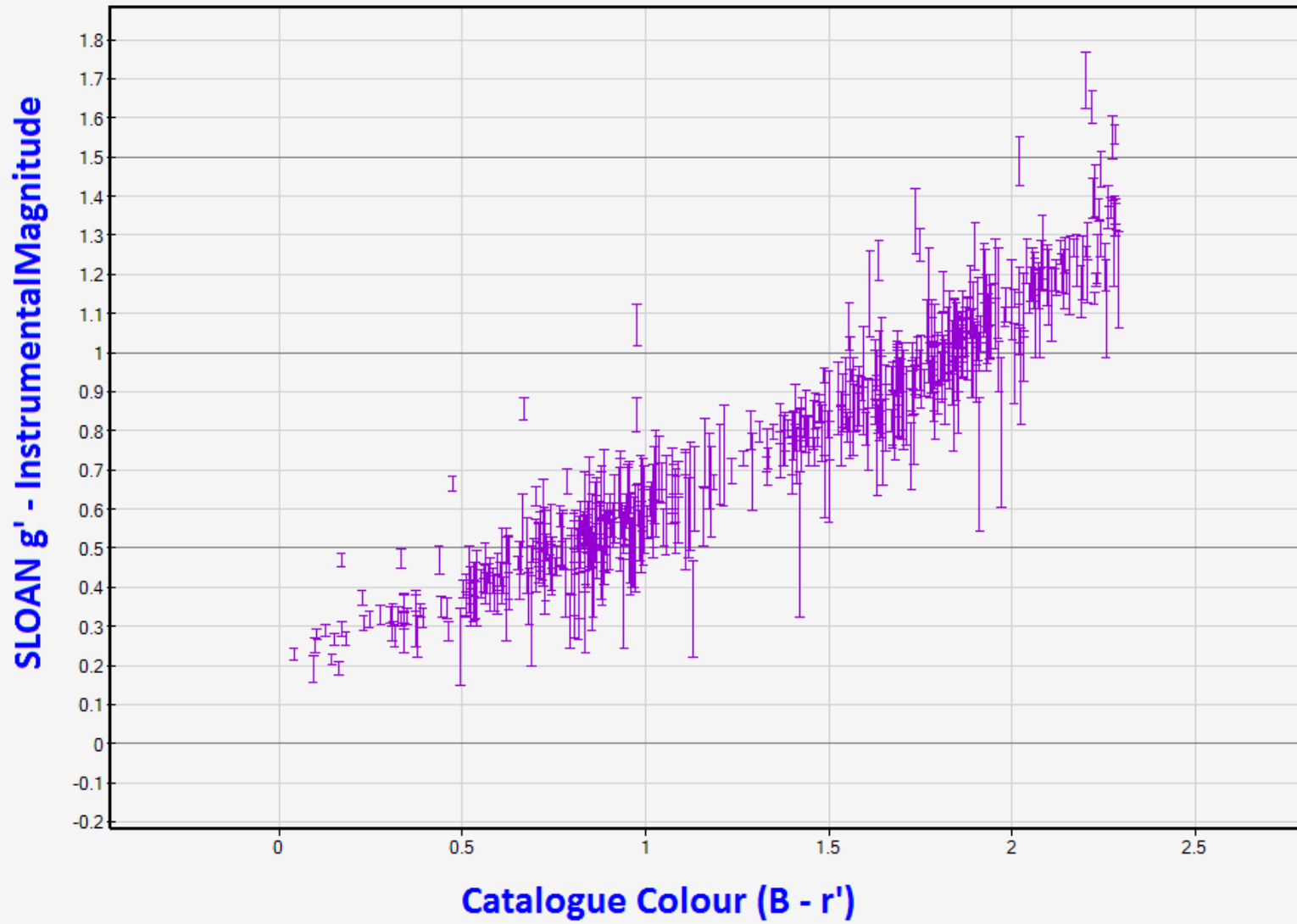
$$r' = -2.5 * 1.138 * \text{Log}_{10} (\text{Intensity}) - 0.041 * (B-r') + 20.32 \pm 0.10$$

### Flea3 (Gamma = OFF)

$$g' = -2.5 * 0.928 * \text{Log}_{10} (\text{Intensity}) + 0.568 * (B-r') + 21.38 \pm 0.14$$

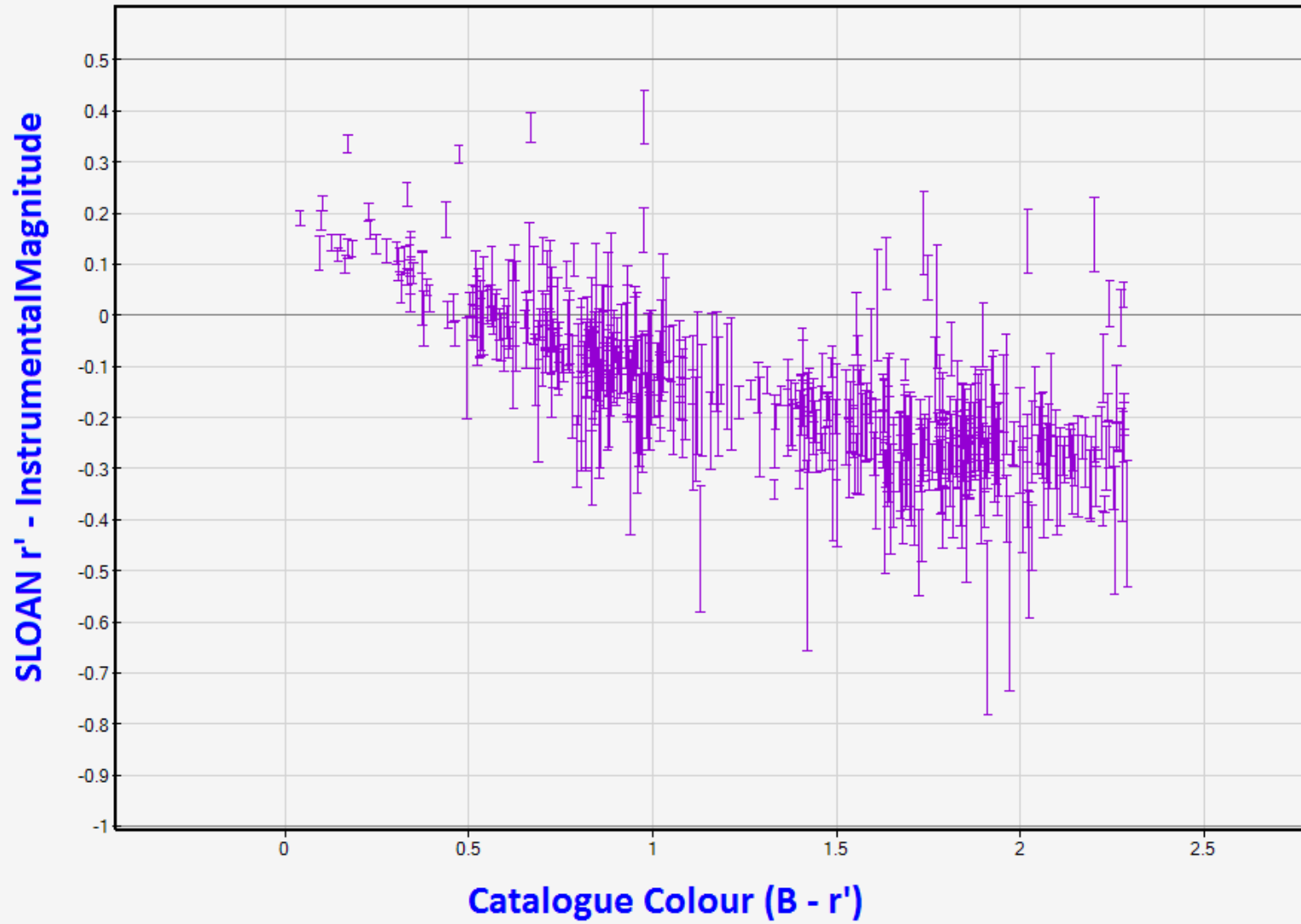
$$r' = -2.5 * 0.927 * \text{Log}_{10} (\text{Intensity}) - 0.102 * (B-r') + 21.49 \pm 0.14$$

# Flea3

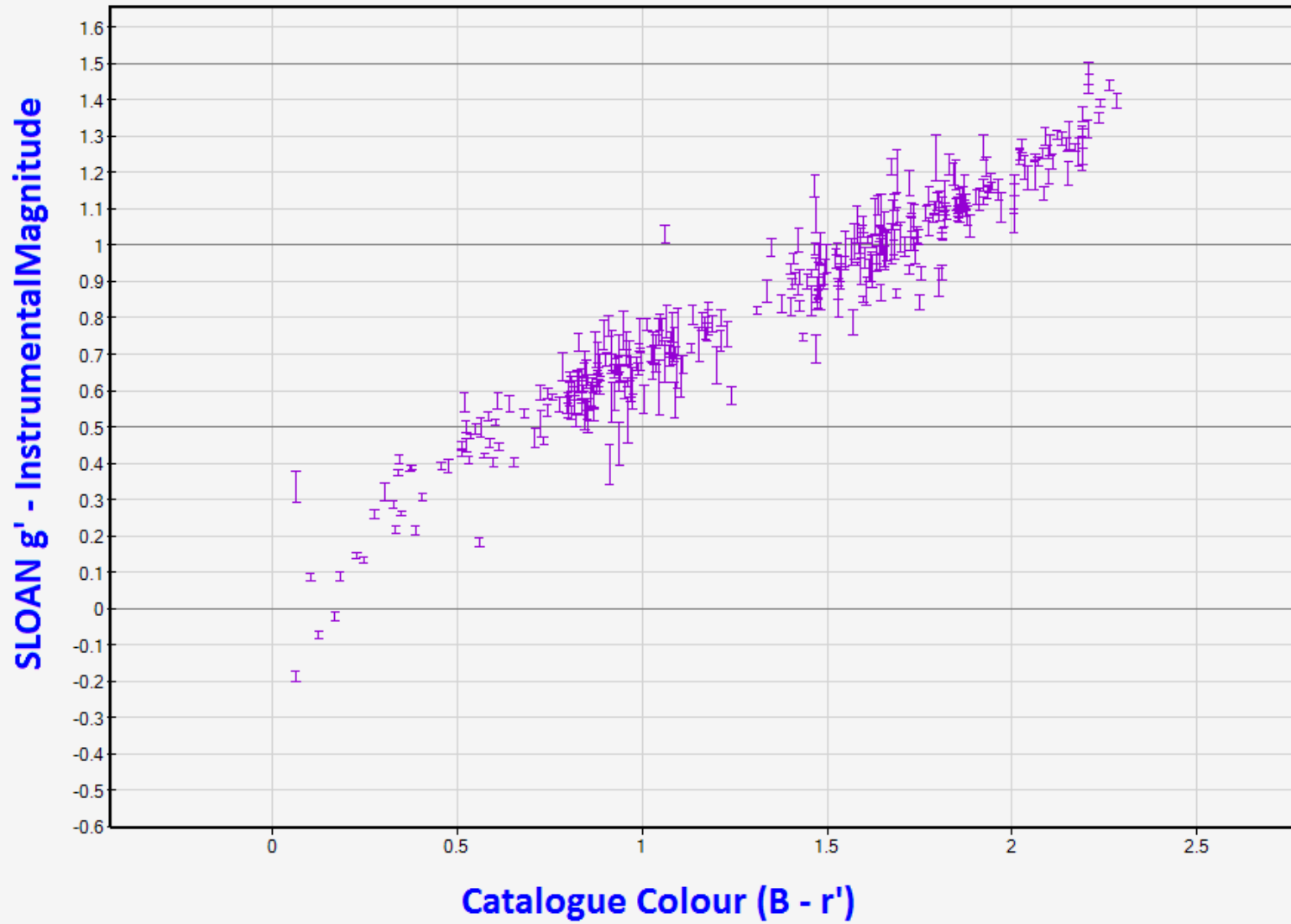




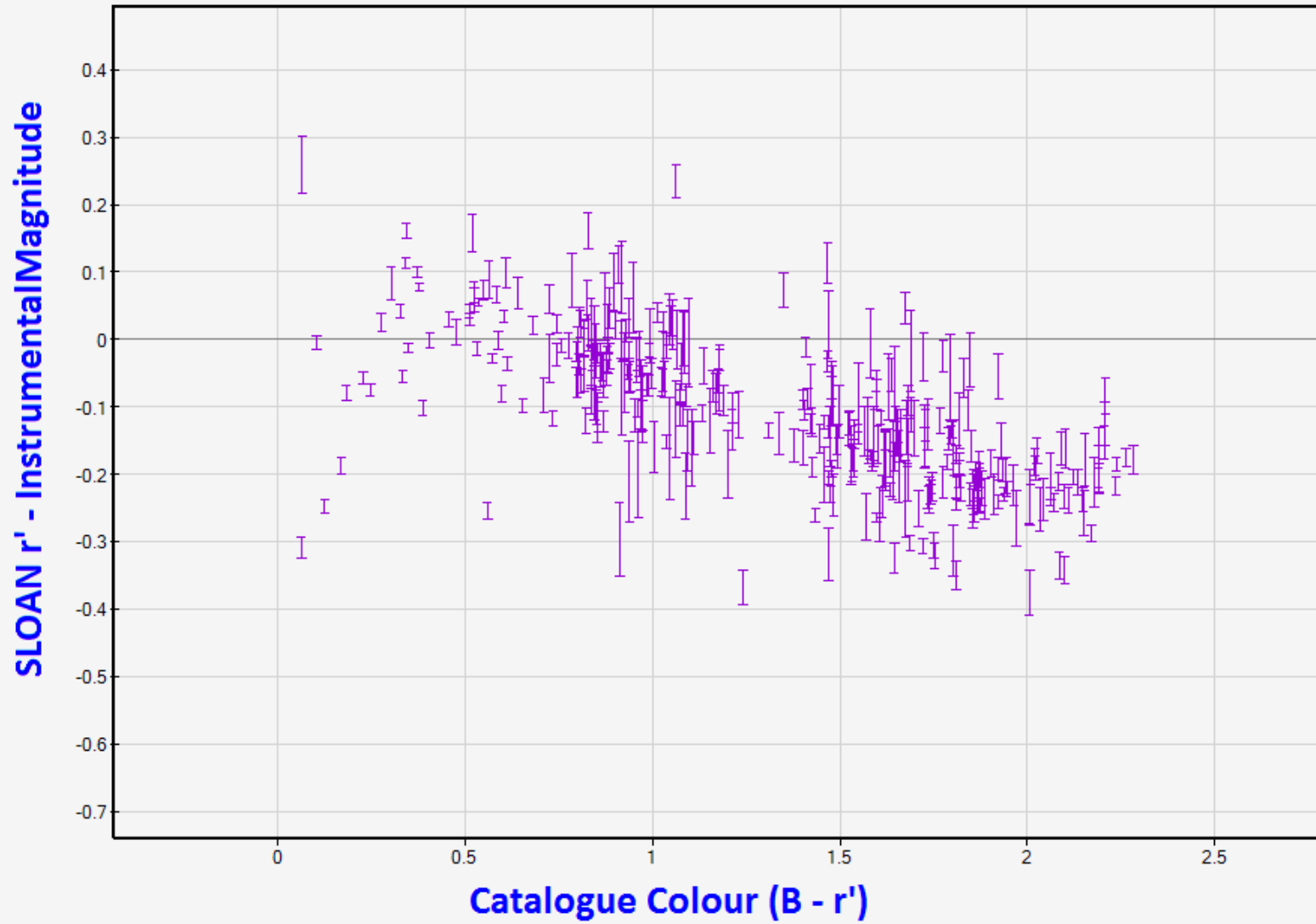
# Flea3



# WAT120N+



# WAT120N+



## WAT120N+

$$g' - \text{InstrumentalMag} = + 0.792 * (B-V) + \text{ZeroPoint} +/- 0.09$$

$$r' - \text{InstrumentalMag} = - 0.169 * (B-V) + \text{ZeroPoint} +/- 0.09$$

$$V - \text{InstrumentalMag} = + 0.271 * (B-V) + \text{ZeroPoint} +/- 0.09$$

$$U3 - \text{InstrumentalMag} = - 0.085 * (B-V) + \text{ZeroPoint} +/- 0.13$$

## WAT902H

$$g' - \text{InstrumentalMag} = + 0.908 * (B-V) + \text{ZeroPoint} +/- 0.13$$

$$r' - \text{InstrumentalMag} = - 0.068 * (B-V) + \text{ZeroPoint} +/- 0.12$$

$$V - \text{InstrumentalMag} = + 0.387 * (B-V) + \text{ZeroPoint} +/- 0.14$$

$$U3 - \text{InstrumentalMag} = + 0.021 * (B-V) + \text{ZeroPoint} +/- 0.19$$

## PC164EX2

$$g' - \text{InstrumentalMag} = + 0.901 * (B-V) + \text{ZeroPoint} +/- 0.12$$

$$r' - \text{InstrumentalMag} = - 0.003 * (B-V) + \text{ZeroPoint} +/- 0.13$$

$$V - \text{InstrumentalMag} = + 0.443 * (B-V) + \text{ZeroPoint} +/- 0.14$$

$$U3 - \text{InstrumentalMag} = + 0.105 * (B-V) + \text{ZeroPoint} +/- 0.21$$

## Flea3 (ICX414) ( for B-V < 1.7)

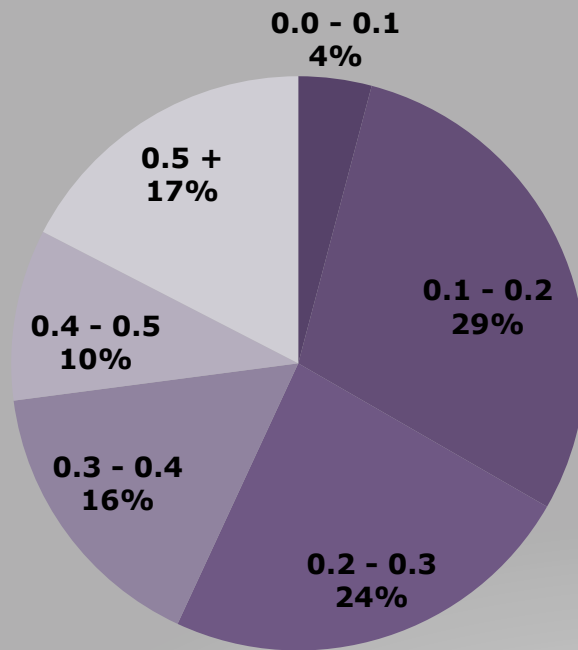
$$g' - \text{InstrumentalMag} = + 0.715 * (B-V) + \text{ZeroPoint} +/- 0.10$$

$$r' - \text{InstrumentalMag} = - 0.256 * (B-V) + \text{ZeroPoint} +/- 0.10$$

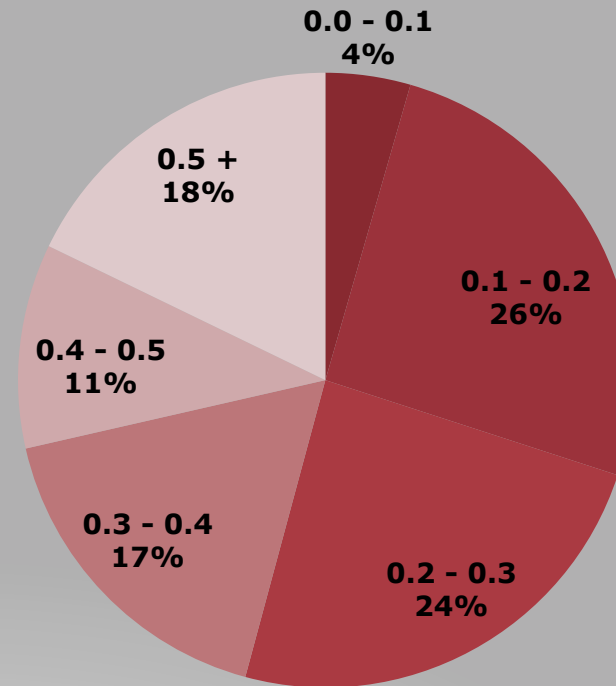
$$V - \text{InstrumentalMag} = + 0.188 * (B-V) + \text{ZeroPoint} +/- 0.10$$

$$U3 - \text{InstrumentalMag} = - 0.156 * (B-V) + \text{ZeroPoint} +/- 0.13$$

### Magn Drop Difference (V)

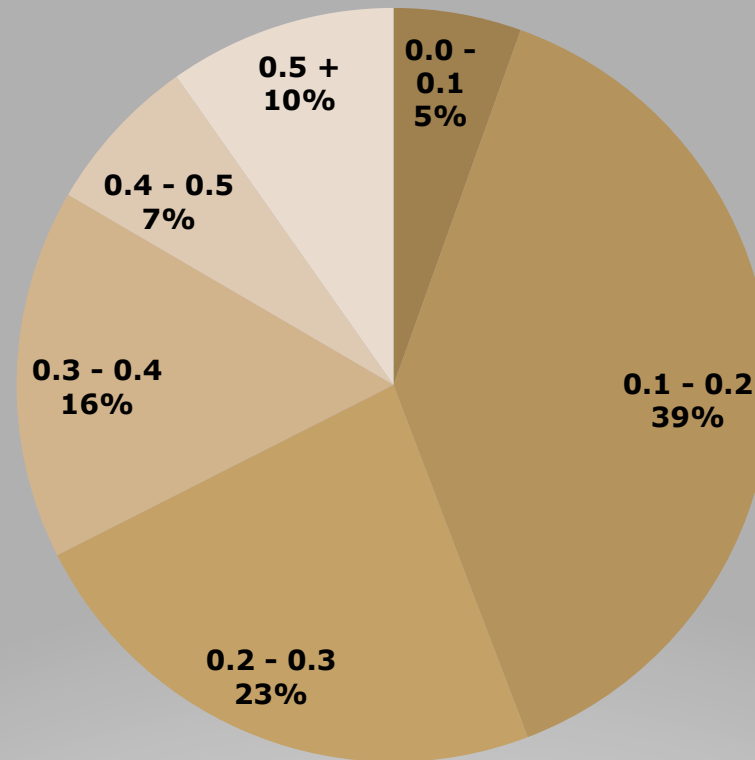


### Magn Drop Difference (R or V where R not available)



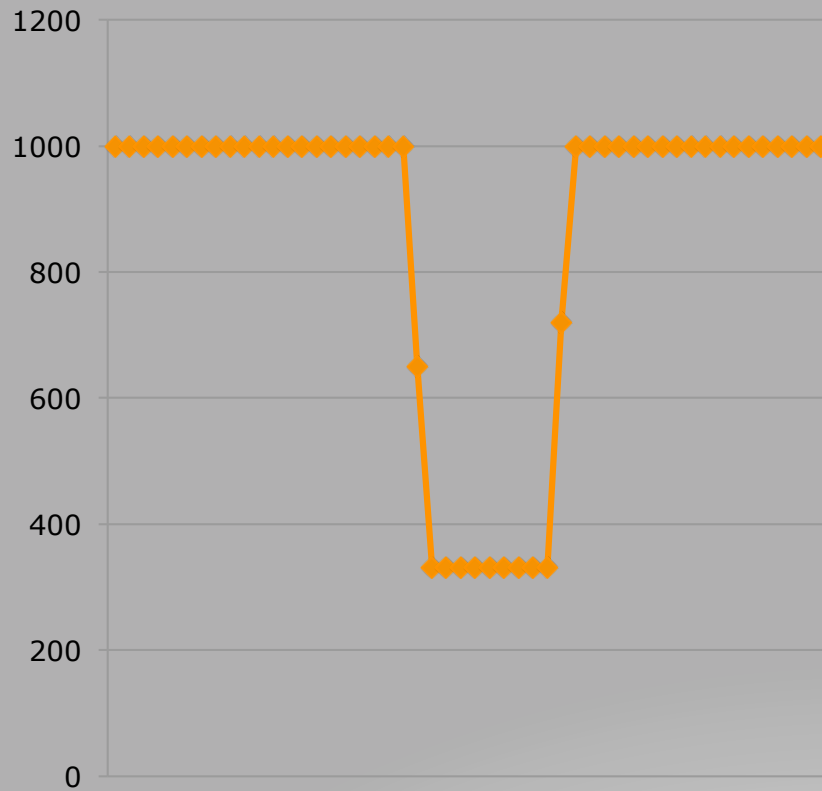
**Steve Preston Events 2012-13**

## Magn Drop Difference (R or V where R not available)



**Partial Mag Drops Only –  
Asteroid Brighter than Mag 13**

### Predicted Drop (1.28<sup>m</sup>)

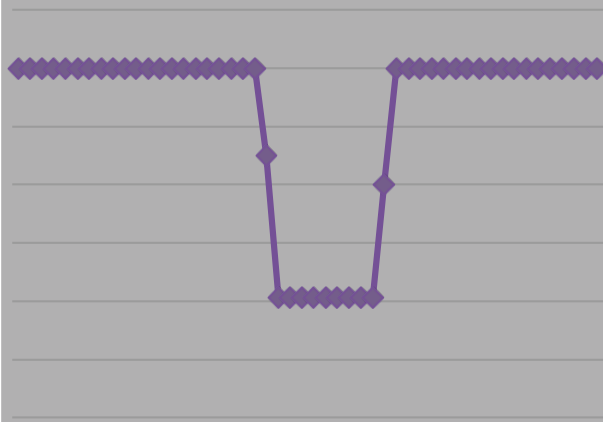


### Observed Drop (0.69<sup>m</sup>)

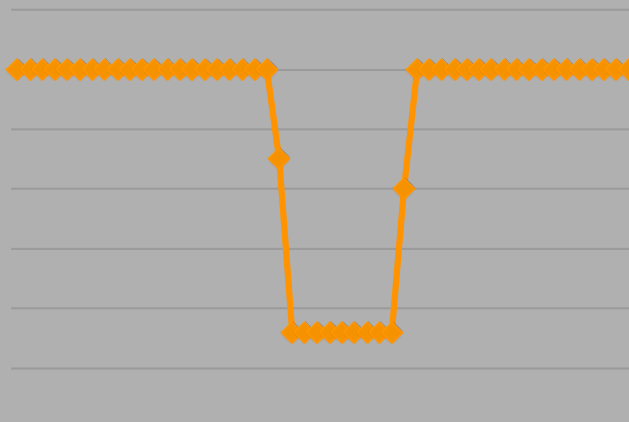


(774) Armor Occults TYC 0567-00326-1  
On 8 Aug 2013

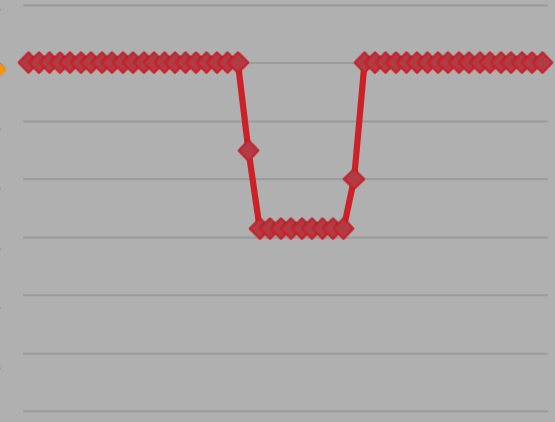
**Predicted (0.54<sup>m</sup>)**



**PC164EX2 (0.62<sup>m</sup>)**



**Flea3 (0.35<sup>m</sup>)**



(1) Ceres Occults TYC 1251-00358-1  
On 04 Jun 2012



- Camera instrumental magnitudes can be quite different than predicted V and R magnitudes
- To match camera measured magnitudes to V or R calibration is required
- Transformation colour coefficients can be determined experimentally for your camera and used to convert standard magnitudes to instrumental magnitudes and vice versa

## Conclusions

# QUESTIONS?

<http://www.hristopavlov.net/videocameras>